

Relativistic gases

$$E^2 = p^2 c^2 + m^2 c^4$$

Ultrarelativistic limit

$$p^2 c^2 \gg m^2 c^4 \Leftrightarrow p \gg mc \quad (\text{muons, neutrinos})$$

$$\Rightarrow E \approx pc = \hbar kc$$

Classical ideal gas

$$Z_1 = \sum_s e^{-\beta E_s}$$

$$g(k)dk = \frac{V k^2}{2\pi^2} dk \Leftrightarrow g(p)dp = \frac{4\pi V p^2}{h^3} dp$$

$$Z_1 = \int_0^\infty g(p) e^{-\frac{\beta p^2}{2m}} dp = V \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2}$$

Ultrarelativistic ideal gas

$$Z_1 = \sum_s e^{-\beta E_s}$$

$$E = pc = \hbar kc$$

$$g(k)dk = \frac{V k^2}{2\pi^2} dk \Leftrightarrow g(p)dp = \frac{V p^2}{2\pi^2 \hbar^3} dp$$

$$Z_1 = \int_0^\infty g(p) e^{-\beta pc} dp = \int_0^\infty \frac{V p^2}{2\pi^2 \hbar^3} e^{-\beta pc} dp = \frac{V}{2\pi^2 (\hbar c)^3} \int_0^\infty x^2 e^{-x} dx$$

$$\left[\int_0^\infty x^2 e^{-x} dx = \left(\frac{d^2}{dx^2} \int_0^\infty e^{-x} dx \right)_{x=1} = \frac{d^2}{dx^2} \left(\frac{1}{x} \right)_{x=1} = \frac{d}{dx} \left(-\frac{1}{x^2} \right)_{x=1} = \frac{2}{x^3} \Big|_{x=1} = 2! \right]$$

$$Z_1 = \frac{V}{2\pi^2 (\hbar c)^3} \cdot 2! = \frac{V}{\pi^2} \left(\frac{k_B T}{\hbar c} \right)^3 = \frac{V}{\lambda^3}$$

$$Z_N = \frac{1}{N!} Z_1^N = \frac{V^N}{N!} \left(\frac{k_B T}{\hbar c} \right)^{3N} = \frac{V^N}{N!} \left(\frac{1}{\lambda} \right)^{3N}$$

thermal wavelength $\lambda = \frac{h}{\sqrt{2\pi m k_B T}} \propto T^{-1/2}$

Helmholtz free energy of UR gas

$$F = -k_B T \ln(Z_N) = -k_B T [N \ln(V) - N \ln(N) + N - 3N \ln(\lambda)] =$$

$$= -k_B T [N \ln\left(\frac{V}{N}\right) + N - 3N \ln(\lambda)] = -N k_B T [-\ln(n) - 3 \ln(\lambda)]$$

$$= N k_B T [\ln(n) + 3 \ln(\lambda) - 1] = N k_B T [\ln(n \lambda^3) - 1]$$

$$F = N k_B T [\ln(n \lambda^3) - 1]$$

Pressure of UR gas

$$p = -\left(\frac{\partial F}{\partial V} \right)_T = -\frac{\partial}{\partial V} (-k_B T N \ln V) = \frac{N k_B T}{V} \Rightarrow pV = N k_B T$$

indep. of β

$$\beta F = N [\ln(h) + 3 \ln \beta + \ln \Lambda^3 - 1]$$

Energy of UR gas

$$U = \frac{\partial \beta F}{\partial \beta} = \frac{\partial}{\partial \beta} (3N \ln \beta) = \frac{3N}{\beta}$$

$$\Rightarrow U = 3Nk_B T$$

$$C_V = \frac{\partial U}{\partial T} = 3Nk_B$$

Comparison

<u>UR</u>	vs.	<u>classical</u>
$p = \frac{Nk_B T}{V}$		
$U = 3Nk_B T$	\neq	$U = \frac{3}{2} Nk_B T$
$C_V = 3Nk_B$	\neq	$C_V = \frac{5}{2} Nk_B$
$p = \frac{1}{3} \frac{U}{V} = \frac{1}{3} u$	\neq	$p = \frac{2}{3} \frac{U}{V} = \frac{2}{3} u$

} 1/2 x

} 2x

Entropy

$$S = \frac{U - F}{T} = 3Nk_B - Nk_B [\ln(h \Lambda^3) - 1] = Nk_B [4 - \ln(h \Lambda^3)]$$

$$S = Nk_B [4 - \ln(h \Lambda^3)]$$

adiabatic expansion of UR gas

$$dS = 0 \Rightarrow S = \text{const.}$$

$$h \Lambda^3 = \text{const.} \Rightarrow \frac{N}{V} \left(\frac{h^2}{2\pi m k_B T} \right)^{3/2} = \text{const.} \Rightarrow VT^3 \text{ is const.}$$

Radiation and matter in the universe

$\rho = \frac{c\rho}{cV}$

$PV^{4/3} = \text{const.}$

$PV^{1/3} = \text{const.}$

\hookrightarrow CMB γ as UR gas in an adiabatically expanding universe

relativistic $\Rightarrow VT^3 = \text{const.} \Rightarrow T \propto V^{-1/3} \propto \frac{1}{a}$

non-relativistic $\Rightarrow VT^{3/2} = \text{const.} \Rightarrow T \propto V^{-2/3} \propto \frac{1}{a^2}$

non-relativistic gas
cools faster than relativistic

relativistic

 $\rho \propto \frac{U}{a^3} \propto \rho \propto V^{-4/3} \propto \frac{1}{a^4}$

$\rho = \frac{1}{3} u$

$\rho \propto \frac{1}{a^4}$

photons

non-relativistic

 $\rho \propto \frac{1}{V} \propto \frac{1}{a^3}$

$\rho \propto \frac{1}{a^3}$

matter

\hookrightarrow universe changed from radiation dominated to matter dominated during expansion

\nearrow after Big Bang

Properties of real gases

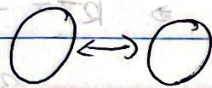
Ideal gas $pV = nRT$

- dilute gas - negligible interaction between gas particles
- gas particles don't take up space

But in real:



part of volume is excluded by gas particles

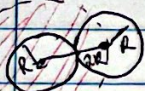


attractive forces between gas particles

Van der Waals

$$V_{\text{excluded}} = \frac{1}{2} n N_A \frac{4}{3} \pi (2R)^3 = n N_A \frac{16}{3} R^3 = nb$$

$$p(V-nb) = nRT \Rightarrow p = \frac{nRT}{V-nb}$$



- ~~pressure is reduced~~ attractive forces reduce pressure on wall
- # particles is influenced $\propto \frac{nN_A}{V}$
- pressure reduction $\propto \left(\frac{nN_A}{V}\right)^2 V$ $E_{\text{att}} \propto \frac{nN_A}{V}$

$$p = \frac{nRT}{V-nb} - a \frac{n^2}{V^2} \Rightarrow \left(p + a \frac{n^2}{V^2}\right)(V-nb) = nRT$$

parameter for prop to

• define molar volume $V_m = \frac{V}{n}$

$$\left(p + \frac{a}{V_m^2}\right)(V_m - b) = RT$$

$V_{\text{excluded}} = n \frac{4}{3} \pi (2R)^3 = nb$ because actual volume is smaller

van der Waals radius $R = \frac{d}{2}$

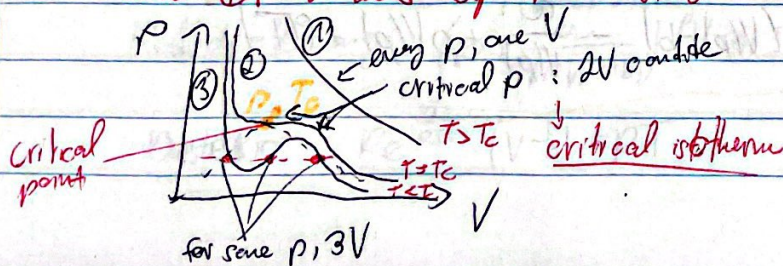
$$b_{\text{argon}} = 0.03219 \cdot 10^{-3} \Rightarrow R = 1.47 \cdot 10^{-10} \text{ m}$$

covalent radius : $R_{\text{argon}} = 0.71 \cdot 10^{-10} \text{ m}$ ← squished together

Van der Waals equation of state:

$$pV^3 - (pb + RT)V^2 + aV - ab = 0$$

cubic in $V \Rightarrow 1, 2, 3$ roots



Critical point

$$\left(p + \frac{a}{V_m^2}\right)(V_m - b) = RT \Rightarrow p = \frac{RT}{V_m - b} - \frac{a}{V_m^2}$$

↳ critical point is inflection point

$$\left(\frac{\partial p}{\partial V_m}\right)_T = 0, \quad \left(\frac{\partial^2 p}{\partial V_m^2}\right)_T = 0, \quad p = \frac{RT}{V_m - b} - \frac{a}{V_m^2}$$

$$\left(\frac{\partial p}{\partial V_m}\right)_T = -\frac{RT}{(V_m - b)^2} + \frac{2a}{V_m^3} = 0 \Rightarrow RT = \frac{2a(V_m - b)^2}{V_m^3}$$

$$\left(\frac{\partial^2 p}{\partial V_m^2}\right)_T = \frac{2RT}{(V_m - b)^3} - \frac{6a}{V_m^4} = 0 \Rightarrow RT = \frac{3a(V_m - b)^3}{V_m^4}$$

$$\Rightarrow 3 \frac{V_m - b}{V_m} = 2$$

$$\downarrow$$

$$V_m^c = 3b$$

$$RT = \frac{2a(3b - b)^2}{27b^3} = \frac{2a \cdot 4b^2}{27b^3} = \frac{8a}{27b}$$

$$\Rightarrow T_c = \frac{8a}{27Rb}$$

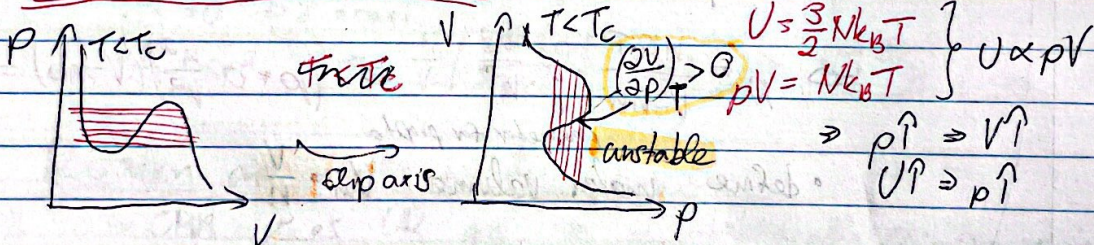
$$p_c = \frac{a}{27b^2}$$

• compression factor Van der Waals gas

$$\frac{p_c V_c}{RT_c} = \left(\frac{3}{8}\right)^n$$

(vs. ideal gas: $\frac{pV}{RT} = 1$)

Van der Waals isotherm $T < T_c$



Gibbs function of VdW gas

$$p = -\left(\frac{\partial F}{\partial V_m}\right)_T \Rightarrow F = -\int p dV_m + f(T) = -\int \left(\frac{RT}{V_m - b} - \frac{a}{V_m^2}\right) dV_m + f(T)$$

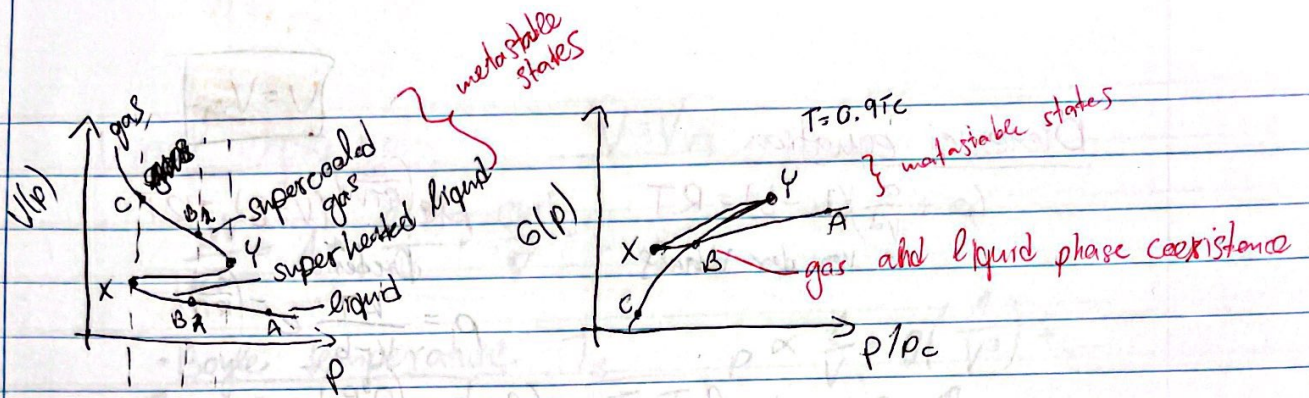
$$F = -RT \ln(V_m - b) - \frac{a}{V_m} + f(T)$$

$$G = F + pV_m = -RT \ln(V_m - b) - \frac{a}{V_m} + pV_m + f(T)$$

$$pV_m^3 - (pb + RT)V_m^2 + aV_m - ab = 0$$

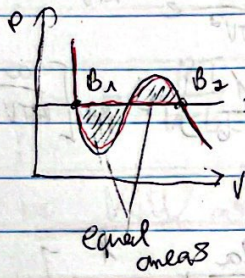
$V(p)$ - not always single valued

$$G(p) = -RT \ln(V(p) - b) - \frac{a}{V(p)} + pV(p) + f(T)$$



Maxwell construction

$$G(p_{B_2}) - G(p_{B_1}) = \int_{p_{B_1}}^{p_{B_2}} \left(\frac{\partial G}{\partial p} \right)_T dp = 0 \text{ along isotherm}$$



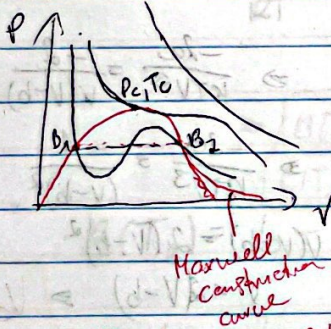
$$G = U - TS + pV$$

$$dG = dU - TdS - SdT + pdV + Vdp$$

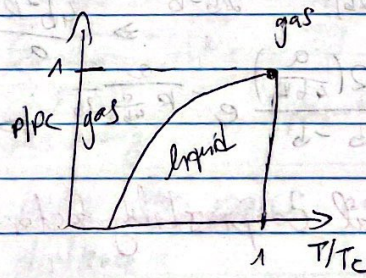
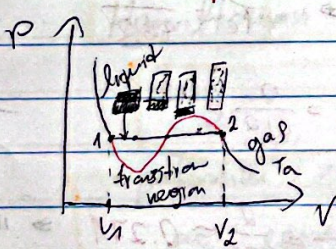
$$dG = -SdT + Vdp \Rightarrow \boxed{V = \left(\frac{\partial G}{\partial p} \right)_T}$$

$$G(p_{B_2}) - G(p_{B_1}) = \int_{p_{B_1}}^{p_{B_2}} V dp = 0 \text{ along isotherm}$$

? How to find B_1 & B_2



Phase equilibrium



Equations of state [for 1 mole]

Van der Waals $\left(p + \frac{a}{V^2} \right) (V - b) = RT$

Berthelot $\left(p + \frac{a}{TV^2} \right) (V - b) = RT$

Clausius $\left(p + \frac{a}{T(V-c)^2} \right) (V - b) = RT$

Redlich-Kwong $\left(p + \frac{a}{\sqrt{T}V(V+b)} \right) (V - b) = RT$

Dieterici $p e^{\frac{a}{RTV}} (V - b) = RT$

$$V = V_m$$

Dieterici equation

$$\left(p + \frac{a}{V^2}\right)(V-b) = RT$$

von der Waals

$$pe^{\left(\frac{a}{RTV}\right)}(V-b) = RT$$

Dieterici

$$p = \frac{RT}{V-b} e^{-\left(\frac{a}{RTV}\right)}$$

• find critical point $\left(\frac{\partial p}{\partial V}\right)_T = \left(\frac{\partial^2 p}{\partial V^2}\right)_T = 0$
 ↳ inflection

$$\left(\frac{\partial p}{\partial V}\right)_T = -\frac{RT}{(V-b)^2} e^{-\frac{a}{RTV}} + \frac{RT}{V-b} \frac{a}{RTV^2} e^{-\frac{a}{RTV}}$$

$$= \frac{RT}{V-b} e^{-\frac{a}{RTV}} \left[\frac{a}{RTV^2} - \frac{1}{V-b} \right] = p \left[\frac{a}{RTV^2} - \frac{1}{V-b} \right] = 0$$

$$\left(\frac{\partial^2 p}{\partial V^2}\right)_T = \left(\frac{\partial p}{\partial V}\right)_T \left[\frac{a}{RTV^2} - \frac{1}{V-b} \right] + p \left[\frac{-2a}{RTV^3} + \frac{1}{(V-b)^2} \right] = 0$$

= 0 at inflection

$$\Rightarrow \frac{a}{RTV^2} - \frac{1}{V-b} = 0 \quad (V, p = 0) \quad \Rightarrow \frac{-2a}{RTV^3} = \frac{-2}{V(V-b)}$$

$$\frac{-2a}{RTV^3} + \frac{1}{(V-b)^2} = 0 \quad \Rightarrow \frac{-2a}{RTV^3} = \frac{-1}{(V-b)^2}$$

$$\Rightarrow V(V-b) = 2(V-b)^2$$

$$V = 2(V-b) \Rightarrow V_c = 2b$$

$$\frac{a}{RT(2b)^2} - \frac{1}{2b-b} = 0 \Rightarrow \frac{4b^2 RT}{a} = b$$

$$T_c = \frac{a}{4bR}$$

$$p_c = \frac{R \left(\frac{a}{4bR}\right)}{2b-b} e^{-\frac{a}{R \frac{a}{4bR} 2b}} \Rightarrow p_c = \frac{a}{4b^2} e^{-2}$$

a, b differs per gas
 ⇒ differs per gas

critical compressibility factor:

$$\frac{p_c V_c}{RT_c} = \frac{2}{e^2} \approx 0.271 \Rightarrow \text{indep. of type of gas}$$

vs. Van der Waals $\frac{p_c V_c}{RT_c} = \frac{3}{8} = 0.375$ ← better

Virial expansion

$$V = V_m$$

$$\frac{pV}{RT} = 1 + \frac{B(T)}{V} + \frac{C(T)}{V^2} + \dots$$

• Boyle temperature T_B $p \propto \frac{1}{V} + O\left(\frac{1}{V^3}\right) + \dots$
 $\hookrightarrow B(T_B) \equiv 0 \Rightarrow T = T_B$

• Boyle's law $p \propto \frac{1}{V}$ for const. V \rightarrow small

$$p = \frac{RT}{V-b} - \frac{a}{V^2} = \frac{RT}{V} \left(\frac{1}{1-\frac{b}{V}} \right) - \frac{a}{V^2}$$

geometric series

for VdW gas

$$\frac{pV}{RT} = \left(1 + \frac{b}{V} + \left(\frac{b}{V}\right)^2 + \dots \right) - \frac{a}{RTV}$$

$$\frac{pV}{RT} = 1 + \frac{b - \frac{a}{RT}}{V} + \left(\frac{b}{V}\right)^2 + \dots$$

$B(T)$ $C(T)$

compare with:

$$\frac{pV}{RT} = 1 + \frac{B(T)}{V} + \frac{C(T)}{V^2} + \dots$$

$$\Rightarrow \begin{cases} B(T) = b - \frac{a}{RT} \\ C(T) = b^2 \end{cases}$$

VdW similar to ideal gas

$$B(T_B) = b - \frac{a}{RT_B} = 0 \Rightarrow T_B = \frac{a}{bR} = \frac{27}{8} T_G$$

for N particle gas

$$\text{Hamiltonian: } H = \underbrace{\sum_{i=1}^N \frac{p_i^2}{2m}}_K + \underbrace{U(\vec{x}_1, \dots, \vec{x}_N)}_{\text{potential energy}}$$

• potential energy only for pairs and only depends on separation distance:

$$U(\vec{x}_1, \dots, \vec{x}_N) = \sum_{i < j} V(|\vec{x}_i - \vec{x}_j|)$$

Partition function

$$Z = \underbrace{\int e^{-\beta \sum_{i=1}^N \frac{p_i^2}{2m}} d\vec{p}_1 \dots d\vec{p}_N}_{Z_{kin}} \underbrace{\int e^{-\beta U(\vec{x}_1, \dots, \vec{x}_N)} d\vec{x}_1 \dots d\vec{x}_N}_{Z_{pot}}$$

$$Z \approx Z_{kin} \cdot Z_{pot} = e^{\dots} \int e^{-\beta U} d\vec{x}_1 \dots d\vec{x}_N$$

ideal gas law

ideal gas
↓

$$Z = Z_{kin} \cdot Z_{pot} \xrightarrow{U=0} Z_{kin}$$

$$Z_{pot}(U=0) = 1 = C \int d\vec{x}_1 \dots d\vec{x}_N = C \cdot V^N \Rightarrow C = \frac{1}{V^N}$$

$$\Rightarrow Z_{pot} = \frac{1}{V^N} \int e^{-\beta U} d\vec{x}_1 \dots d\vec{x}_N$$

$$U = \sum_{i < j} V(|\vec{x}_i - \vec{x}_j|)$$

to prepare to approx.
for $x \ll \lambda$

$$Z_{pot} = \frac{1}{V^N} \int e^{-\beta U} d\vec{x}_1 \dots d\vec{x}_N$$

$$Z_{pot} = 1 + \frac{1}{V^N} \int (e^{-\beta \sum_{i < j} V(|\vec{x}_i - \vec{x}_j|)} - 1) d\vec{x}_1 \dots d\vec{x}_N$$

only important when particles close together ($|\vec{x}_i - \vec{x}_j|$ small)

↳ suppose gas is dilute and close to ideal

⇒ only two particles close together ⇒ only 1 pair at a time interacting

$$\frac{N(N-1)}{2} \approx \frac{N^2}{2} \text{ different pairs}$$

$$Z_{pot} = 1 + \frac{N^2}{2V^N} \int (e^{-\beta V(|\vec{x}_1 - \vec{x}_2|)} - 1) d\vec{x}_1 \dots d\vec{x}_N$$

$$Z_{pot} = 1 + \frac{N^2}{2V^N} V^{N-2} \int (e^{-\beta V(|\vec{x}_1 - \vec{x}_2|)} - 1) d\vec{x}_1 d\vec{x}_2$$

we can integrate over $\int d\vec{x}_3 \dots d\vec{x}_N = V^{N-2}$

only relative distance matters

$$Z_{pot} = 1 + \frac{N^2}{2V^2} V \int (e^{-\beta V(|\vec{x}|)} - 1) d\vec{x}$$

integrate one more coord. vector pointing from one particle to the other

$$Z_{pot} = 1 + \frac{N^2}{2V^2} \int (e^{-\beta V(|\vec{x}|)} - 1) d\vec{x} = 1 - \frac{N}{V} B'(T)$$

$$Z_{pot} = 1 - \frac{N}{V} B'(T)$$

Thermodynamics

$$F = -k_B T \ln Z = -k_B T \ln(Z_{kin} Z_{pot}) =$$

$$= -k_B T \ln(Z_{kin}) - k_B T \ln(Z_{pot}) =$$

$$= F_{kin} - k_B T \ln\left(1 - \frac{N}{V} B'(T)\right)$$

deviations from ideal gas ⇒ assume small

$$\Rightarrow B'(T) \text{ small} \Rightarrow \ln(1+x) \approx x$$

$$-\left(\frac{\partial F}{\partial V}\right)_T$$

$$\approx F_{kin} + \frac{Nk_B T}{V} B'(T)$$

$$p = \frac{Nk_B T}{V} + \frac{Nk_B T}{V^2} B'(T)$$

$$\Rightarrow \frac{pV}{Nk_B T} = 1 + \frac{B'(T)}{V}$$

$$\frac{pV}{Nk_B T} \approx 1 + \frac{B(T)}{V} \Rightarrow \frac{pV}{nN_A k_B T} \approx 1 + \frac{B(T)}{V}$$

$$\Rightarrow \frac{pV_m}{RT} \approx 1 + \frac{B(T)/n}{V_m} = 1 + \frac{B(T)}{V}$$

1st order virial expansion \rightarrow other terms if more particle interaction

Second virial coefficient

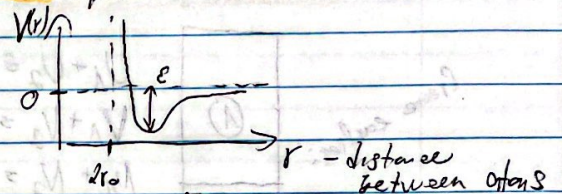
$$B'(T) = \frac{N}{2} \int (1 - e^{-\beta v(r)}) d\vec{x} = \frac{N}{2} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} (1 - e^{-\beta v(x)}) dx dy dz =$$

$$= \frac{N}{2} \int_0^{\infty} \int_0^{2\pi} \int_0^{\pi} (1 - e^{-\beta v(r)}) r^2 \sin\theta d\theta d\phi dr = \frac{N}{2} \int_0^{\infty} 4\pi r^2 (1 - e^{-\beta v(r)}) dr$$

$$\rightarrow B'(T) = \frac{N}{2} \int_0^{\infty} 4\pi r^2 (1 - e^{-\beta v(r)}) dr$$

- potential energy

$$\left. \begin{aligned} v(r) &= \infty & \text{for } r < 2r_0 \\ \beta v(r) &= \frac{v(r)}{k_B T} \ll 1 & \text{for } r > 2r_0 \end{aligned} \right\}$$



$$B'(T) = \frac{N}{2} \int_0^{2r_0} 4\pi r^2 (1 - e^{-\beta v(r)}) dr + \frac{N}{2} \int_{2r_0}^{\infty} 4\pi r^2 (1 - e^{-\beta v(r)}) dr =$$

$$= \frac{N}{2} \int_0^{2r_0} 4\pi r^2 dr + \frac{N}{2} \int_{2r_0}^{\infty} \beta v(r) 4\pi r^2 dr$$

$\approx \beta v$ because $\beta v \ll 1$

$$= \frac{N}{2} 4\pi \frac{1}{3} (2r_0)^3 + \frac{N}{k_B T} 2\pi \int_{2r_0}^{\infty} r^2 v(r) dr =$$

$$= nN_A \left(\frac{16}{3} \pi r_0^3 \right) + \frac{n}{k_B T} 2\pi N_A \int_{2r_0}^{\infty} r v(r) dr = nb - \frac{a}{k_B T}$$

From: $V_{excluded} = nN_A \frac{4}{3} \pi r_0^3 = nb$

$$B'(T) = n \left(b - \frac{a}{k_B T} \right) \Rightarrow B(T) = \frac{B'(T)}{n} = b - \frac{a}{k_B T}$$

$\frac{pV}{RT} \approx 1 + \frac{B(T)}{V}$ Second virial coeff. of van der Waals eq.

Law of corresponding states

using reduced coordinates: $\tilde{p} = \frac{p}{p_c}$, $\tilde{V} = \frac{V}{V_c}$, $\tilde{T} = \frac{T}{T_c}$

$$\left(p + \frac{a}{V^2} \right) (V - b) = RT$$

$$\left(\tilde{p} p_c + \frac{a}{(\tilde{V} V_c)^2} \right) (\tilde{V} V_c - b) = R \tilde{T} T_c$$

$$\tilde{p} + \frac{1}{\tilde{V}^2} \frac{a}{p_c V_c^2} = \frac{\tilde{T}}{\tilde{V} V_c - b} \frac{R T_c}{p_c}$$

$$\bar{p} + \frac{1}{\bar{v}^2} \frac{a}{p_0 \bar{v}_c^2} = \frac{\bar{T}}{\bar{v} \bar{v}_c - b} \frac{RT_c}{p_0} = R \frac{8a}{27Rb} \frac{27b^2}{a} = 8b$$

$$\begin{aligned} V_c &= 3b \\ T_c &= \frac{8a}{27Rb} \\ p_c &= \frac{a}{27b^2} \end{aligned}$$

$$\Rightarrow \left(\bar{p} + \frac{3}{\bar{v}^2} \right) = \frac{8\bar{T}}{3\bar{v} - 1} \Rightarrow \left(\bar{p} + \frac{3}{\bar{v}^2} \right) = \frac{8\bar{T}}{3\bar{v} - 1}$$

compressibility $\Rightarrow \kappa = \frac{pV}{RT} = \frac{\bar{p}\bar{v}}{RT_c} \frac{p_0 \bar{v}_c}{RT_c}$
 $\hookrightarrow 1$ for ideal gas
 $\kappa = \frac{3}{8} \frac{\bar{p}\bar{v}}{\bar{T}} = \frac{3}{8}$ for vdW

b indep. of a and b
 \Rightarrow in reduced coord., the eq. of state is same for every gas

Phase Transitions

Phase equlib:

①	$U_1 + U_2 = U$ $V_1 + V_2 = V$ $N_1 + N_2 = N$
②	$S_1 + S_2 = S$

in equilibrium: $dS = 0$

$$\left(\frac{\partial S}{\partial U} \right)_{V_1, N_1} dU_1 + \left(\frac{\partial S}{\partial V_1} \right)_{U_1, N_1} dV_1 + \left(\frac{\partial S}{\partial N_1} \right)_{U_1, V_1} dN_1 = 0$$

$$\begin{aligned} T_1 &= T_2 \\ p_1 &= p_2 \\ \mu_1 &= \mu_2 \end{aligned} \quad \left(\begin{aligned} p &= T \frac{\partial S}{\partial V} \\ \frac{1}{T} &= \frac{\partial S}{\partial U} \\ \mu &= -T \frac{\partial S}{\partial N} \end{aligned} \right)$$

Equilibrium of two phases

$$dS = \underbrace{\left(\frac{\partial S}{\partial U} \right)}_{= \frac{1}{T}} dU + \underbrace{\left(\frac{\partial S}{\partial V} \right)}_{= \frac{p}{T}} dV + \underbrace{\left(\frac{\partial S}{\partial N} \right)}_{= -\frac{\mu}{T}} dN$$

$$dS = \frac{dU}{T} + \frac{p}{T} dV - \frac{\mu}{T} dN$$

$$dU = TdS - pdV + \mu dN$$

Gibbs free energy

$$G = U + pV - TS$$

$$dG = -SdT + Vdp + \mu dN$$

$$\mu = \left(\frac{\partial G}{\partial N} \right)_{T, p}$$

$$dG = \sum_i \mu_i dN_i$$

\hookrightarrow extensive: $G = N \cdot g$

$$\mu = \left(\frac{\partial G}{\partial N} \right)_{T, p} = g(T, p)$$

Gibbs free energy per particle

alternative way:

equilibrium: $\mu_1 = \mu_2 \Rightarrow g_1(T, p) = g_2(T, p)$

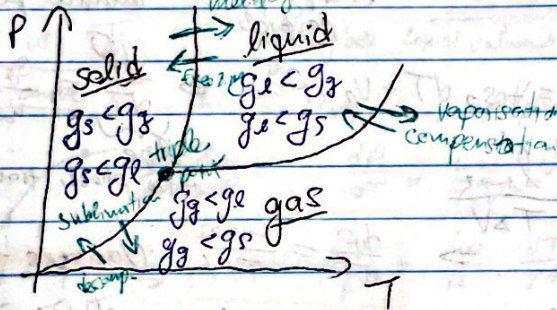
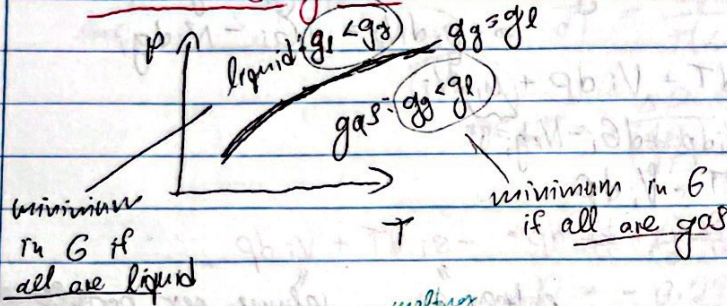
G at minimum $\Rightarrow dG = 0$
and $G = N_1 g_1 + N_2 g_2$

$$dG = \left(N_1 \frac{\partial g_1}{\partial T} + N_2 \frac{\partial g_2}{\partial T} \right) dT + \left(N_1 \frac{\partial g_1}{\partial p} + N_2 \frac{\partial g_2}{\partial p} \right) dp + g_1 dN_1 + g_2 dN_2 = 0$$

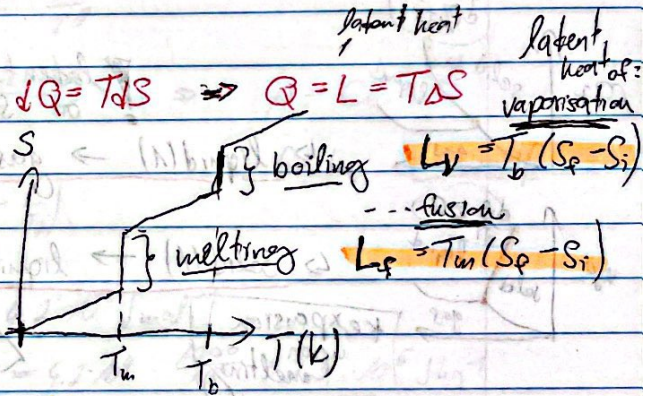
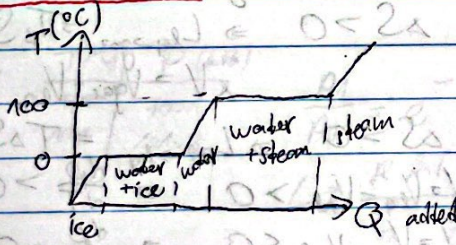
$\begin{matrix} = 0 & & = 0 \\ \text{at equilibrium} & & \text{at equilibrium} \end{matrix}$

$$\Rightarrow (g_1 - g_2) dN_1 = 0 \Rightarrow g_1 = g_2$$

Phase diagrams



Latent heat



Trouton's rule

$$\frac{L_v}{T_b} \approx 10R \approx 83 \text{ kJ/mol}$$

$\approx \Delta S$

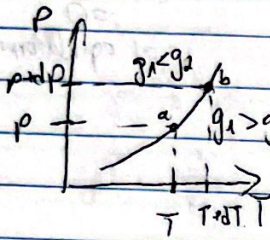
$$\Delta S = \Delta(k \ln \Omega) = k \ln \Omega_v - k \ln \Omega_l = k \ln \left(\frac{\Omega_v}{\Omega_l} \right)$$

$$\frac{\Omega_v}{\Omega_l} \approx \left(\frac{V_v}{V_l} \right)^{N_A} \approx \left(\frac{p_l}{p_v} \right)^{N_A} \approx (10^3)^{N_A}$$

water for example

$$\Rightarrow \Delta S = k \ln (10^3)^{N_A} = \frac{k N_A}{R} \ln(10^3) \approx 7R$$

Clausius - Clapeyron equation



$$g_1(b) = g_2(b)$$

$$g_1(a) = g_2(a)$$

$$\Rightarrow dg_1 = dg_2 \text{ along the phase eq. curve}$$

$$G_i = N_i g_i \Rightarrow dG_i = N_i dg_i + g_i dN_i$$

$$\Rightarrow g_i dN_i = dG_i - N_i dg_i$$

$$\Rightarrow dG_i = -S_i dT + V_i dp + \mu_i dN_i$$

$$dG_i = -S_i dT + V_i dp + dG_i - N_i dg_i$$

$$N_i dg_i = -S_i dT + V_i dp$$

$$\Rightarrow dg_i = -\frac{S_i}{N_i} dT + \frac{V_i}{N_i} dp = -s_i dT + v_i dp$$

entropy per particle

volume per particle

$$dg_1 = dg_2$$

$$-s_1 dT + v_1 dp = -s_2 dT + v_2 dp$$

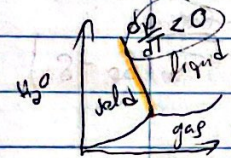
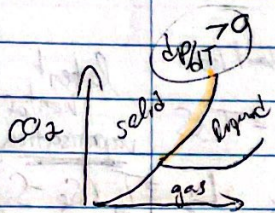
$$\Rightarrow \frac{dp}{dT} = \frac{s_2 - s_1}{v_2 - v_1} \left\{ \begin{array}{l} \text{per particle,} \\ \text{mole, kg, ...} \end{array} \right.$$

slope of the phase transition curve

$$\Rightarrow \frac{dp}{dT} = \frac{\Delta S}{\Delta V} = \frac{L_{1 \rightarrow 2}}{T \Delta V}$$

$$\Rightarrow \frac{dp}{dT} = \frac{L_{1 \rightarrow 2}}{T \Delta V}$$

Clausius - Clapeyron eq.



! latent heat and volume change should be calculated on same basis (mole/kg/...)

$$\hookrightarrow \text{liquid (1)} \rightarrow \text{gas (2)} : \Delta S > 0 \Rightarrow L_{\text{liq} \rightarrow \text{gas}} = T \Delta S > 0$$

$$\Delta V = V_{\text{gas}} - V_{\text{liq}} > 0 \Rightarrow \frac{dp}{dT} > 0$$

$$\hookrightarrow \text{solid (1)} \rightarrow \text{liquid (2)} : \Delta S > 0 \Rightarrow L_{\text{sol} \rightarrow \text{liq}} = T \Delta S > 0$$

expansion upon melting

$$\left\{ \begin{array}{l} \text{if } \Delta V = V_{\text{liq}} - V_{\text{sol}} > 0 \Rightarrow \frac{dp}{dT} > 0 \text{ most substances} \\ \text{if } \Delta V = V_{\text{liq}} - V_{\text{sol}} < 0 \Rightarrow \frac{dp}{dT} < 0 \text{ water, Si, Bi} \end{array} \right.$$

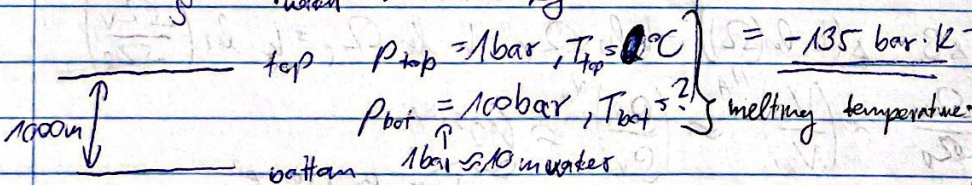
$$\star L_{\text{ice} \rightarrow \text{water}} = 3.35 \cdot 10^5 \text{ J kg}^{-1} \text{ at } 273 \text{ K, } 1 \text{ bar}$$

Glacier

$$\rho^{-1} = V_{\text{ice}} = 1091 \text{ cm}^3/\text{kg}$$

$$\rho^{-1} = V_{\text{water}} = 1000 \text{ cm}^3/\text{kg}$$

$$\frac{dp}{dT} = \frac{L_{\text{ice} \rightarrow \text{water}}}{T(V_{\text{water}} - V_{\text{ice}})} = \frac{3.35 \cdot 10^5}{273.15 \cdot (-91 \cdot 10^{-6})} = -135 \text{ bar} \cdot \text{K}^{-1}$$



$$\frac{\Delta P}{\Delta T} = \frac{P_{\text{bot}} - P_{\text{top}}}{T_{\text{bot}} - T_{\text{top}}} = -135 \text{ bar} \cdot \text{K}^{-1} \Rightarrow \Delta T = \frac{\Delta P}{-135} = \frac{99}{-135} = -0.73 \text{ K}$$

$$T_{\text{bot}} = T_{\text{top}} + \Delta T = -0.73^\circ \text{C}$$

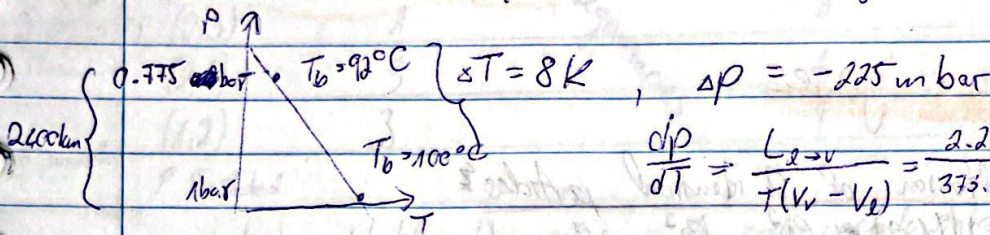
melting \uparrow less than 0°C

when heating up, bottom melts first \Rightarrow glacier slides
 \uparrow
 at lower temp

ff

*** boiling & pressure**

$L_{\text{liquid} \rightarrow \text{vapor}} = 2.26 \cdot 10^6 \text{ J kg}^{-1}$
 $V_v = 1.67 \cdot 10^6 \text{ cm}^3 \text{ kg}^{-1}$



$$\frac{dp}{dT} = \frac{L_{l \rightarrow v}}{T(V_v - V_l)} = \frac{2.26 \cdot 10^6}{373.15 \cdot 1.67} = 3.6 \text{ kPa K}^{-1} = 36 \text{ mbar K}^{-1}$$

$$\Rightarrow \Delta T = \frac{\Delta p}{\left(\frac{dp}{dT}\right)} = \frac{-225}{36} = -6.25 \text{ K}$$

but above: -8 K

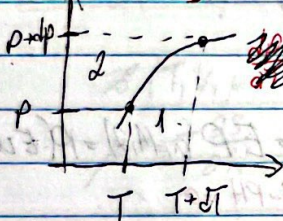
$$P = P_0 e^{-h/\lambda_{km}} = P_0 e^{-2.4/h} = 0.71 P_0$$

$$\Delta p = -0.29 \text{ bar} \Rightarrow \Delta T = -8 \text{ K} \checkmark$$

~~XXXXXXXXXX~~

P vs. T

Vapor pressure curve



negligible liquid volume

$$\Delta V \approx V_{\text{vapor}}, \quad pV_{\text{vapor}} = RT$$

$$\frac{dp}{dT} = \frac{L_{l \rightarrow v}}{T \Delta V} = \frac{L_{l \rightarrow v} p}{T RT} = \frac{L_{l \rightarrow v}}{TR} p$$

$$\frac{dp}{p} = \frac{L_{l \rightarrow v}}{R} \frac{dT}{T^2}$$

$$\ln p = -\frac{L_{l \rightarrow v}}{RT} + C \Rightarrow$$

$$p = C e^{-\frac{L_{l \rightarrow v}}{RT}}$$

for $L_{l \rightarrow v}$ temp. indep.

T (°C)	p (kPa)
90	70.1
100	101.3

$$\frac{p_1}{p_2} = e^{-\frac{L}{R} \left(\frac{1}{T_1} - \frac{1}{T_2} \right)}$$

$$L = \frac{R \ln \left(\frac{p_2}{p_1} \right)}{\frac{1}{T_1} - \frac{1}{T_2}} = 4.2 \cdot 10^4 \text{ J mol}^{-1}$$

$$L = 4.2 \cdot 10^4 \cdot \frac{1000}{18} = 2.3 \cdot 10^6 \text{ J kg}^{-1}$$

L dependant on T

$$dH = T ds + V dp \Big|_{\text{const. } p} = T ds = dQ$$

$$C_p = \left(\frac{\partial Q}{\partial T} \right)_p = \left(\frac{\partial H}{\partial T} \right)_p$$

$$C_{p, \text{vapor}} - C_{p, \text{liq}} = \left(\frac{\partial H_{\text{vapor}}}{\partial T} \right)_p - \left(\frac{\partial H_{\text{liq}}}{\partial T} \right)_p = \left(\frac{\partial (H_{\text{vapor}} - H_{\text{liq}})}{\partial T} \right)_p = \left(\frac{\partial \Delta H}{\partial T} \right)_p$$

$$\Delta G = \Delta H - T \Delta S = \Delta H - L(T) \approx 0 \Rightarrow \left(\frac{\partial L(T)}{\partial T} \right)_p = C_{p, \text{vapor}} - C_{p, \text{liq}}$$

onshore
banking

$$\Rightarrow L(T) = L_0 + (C_{p, \text{vapor}} - C_{p, \text{liq}}) T$$

$$\frac{dp}{dT} = \frac{pL_0(T)}{RT^2} = \frac{p \cdot (L_0 + \Delta CT)}{RT^2}$$

$$\ln p = -\frac{L_0}{RT} + \frac{\Delta C}{R} \ln(T) \Rightarrow$$

$$p = p_0 e^{-\frac{L_0}{RT} + \frac{\Delta C}{R} \ln(T)}$$

Quantum gases

Hamiltonian of 2 identical particles

$$H(1,2) = \frac{p_1^2}{2m} + \frac{p_2^2}{2m} + V(x_1, x_2)$$

$$\Rightarrow H(1,2) = H(2,1) \Rightarrow V(x_1, x_2) = V(x_2, x_1)$$

$$\begin{cases} H(1,2) u_E(1,2) = E u_E(1,2) \\ H(2,1) u_E(2,1) = E u_E(2,1) \end{cases} \text{ identical so we can relabel}$$

$$\Rightarrow \boxed{H(1,2) u_E(2,1) = E u_E(2,1)}$$

to

define exchange operator for 2 particle wavefunction

$$P \psi(1,2) = \psi(2,1)$$

$$\begin{aligned} H(P u_E(1,2)) &= H u_E(2,1) = E u_E(2,1) = E P u_E(1,2) = P(E u_E(1,2)) = \\ &= P(H u_E(1,2)) \Rightarrow [H, P] = HP - PH = 0 \end{aligned}$$

\Rightarrow H and P are commuting operators, thus P is a constant of motion

• eigenvalues of P

$$P^2 \psi(1,2) = P \psi(2,1) = \psi(1,2)$$

$$\Rightarrow P^2 = 1 \Rightarrow \boxed{P = \pm 1}$$

• eigenstates of P

$$P=1: \psi^S(1,2) = \frac{1}{\sqrt{2}} (\psi(1,2) + \psi(2,1))$$

$$P=-1: \psi^A(1,2) = \frac{1}{\sqrt{2}} (\psi(1,2) - \psi(2,1))$$

symmetric - bosons
or
antisymmetric - fermions
under particle exchange

fermions x bosons

antisymmetric

↳ half-integer spin

$$\frac{1}{2}, \frac{3}{2}, \frac{5}{2}$$

↳ symmetric

↳ integer spin

$$0, 1, 2, 3$$

also for compound particles: ${}^3\text{He}$ (spin $\frac{1}{2}$)

${}^4\text{He}$ (spin 0)

• suppose 2 fermions in states \underline{S} and \underline{V} : $\psi^A(1,2) = \frac{1}{\sqrt{2}} (\phi_S(1)\phi_V(2) - \phi_V(1)\phi_S(2))$

- if $\underline{S} = \underline{V}$: $\psi^A(1,2) = 0$

\Rightarrow two fermions cannot be in the same state

Particles and states

1 particle

2 states

3 particles

particles

states

	distinguishable classical	indistinguishable classical & indist. bosons
(3,0)	1	1
(0,3)	1	1
(2,1)	3	1
(1,2)	3	1

degeneracies

indistinguishable fermions

- 3,0
- 0,3
- 2,1
- 1,2

0

- 2 single particle states
3 particles
⇒ not possible - there cannot be more than one identical fermions in same state

Grand partition function - Gibbs distribution

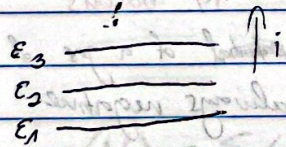
$$P_j = \frac{1}{Z} e^{\beta(\mu N_j - E_j)}$$

$$Z(T, V, \mu) = \sum_j e^{\beta(\mu N_j - E_j)}$$

$$\mu = -T \frac{\partial S}{\partial N}$$

small system

$\{n_1, n_2, n_3, \dots\}$ defines a state j of the small system



$$N_j = \sum_i n_i, \quad E_j = \sum_i n_i E_i$$

$$Z = \sum_{n_1} \sum_{n_2} \dots e^{\beta(\mu \sum_i n_i - \sum_i n_i E_i)} = \sum_{n_1} \sum_{n_2} \dots e^{\beta n_1(\mu - E_1)} e^{\beta n_2(\mu - E_2)} \dots$$

$$= \left(\sum_{n_1} e^{\beta n_1(\mu - E_1)} \right) \left(\sum_{n_2} e^{\beta n_2(\mu - E_2)} \right) \dots$$

$$\Rightarrow Z = \prod_{i=1}^{\infty} Z_i, \quad Z_i = \sum_{n_i} e^{\beta n_i(\mu - E_i)}$$

probability that the gas is in a state: $\{n_1, n_2, n_3, \dots\}$

$$P(n_1, n_2, \dots) = \frac{e^{\beta(\mu(n_1+n_2+\dots) - (n_1 E_1 + n_2 E_2 + \dots))}}{Z} = \frac{e^{\beta n_1(\mu - E_1)}}{Z_1} \frac{e^{\beta n_2(\mu - E_2)}}{Z_2} \dots$$

normalised

$$\Rightarrow P(n_1, n_2, \dots) = \prod_{i=1}^{\infty} P_i(n_i), \quad P_i(n_i) = \frac{e^{\beta n_i(\mu - E_i)}}{Z_i}$$

probability of finding n_i particles in single particle state i is indep. of # particles in other states

Mean occupation numbers

$$\langle n_i \rangle = \sum_{n_1, n_2, \dots} n_i P(n_1, n_2, \dots) = \sum_{n_1, n_2, \dots} n_i P_1(n_1) P_2(n_2) \dots P_i(n_i) \dots$$

$$= \underbrace{\left(\sum_{n_1} P_1(n_1) \right)}_{=1} \underbrace{\left(\sum_{n_2} P_2(n_2) \right)}_{=1} \dots \left(\sum_{n_i} n_i P_i(n_i) \right)$$

$$\langle n_i \rangle = \sum_{n_i} n_i P_i(n_i) = \sum_{n_i} n_i e^{\beta n_i (\mu - \epsilon_i)} = \frac{1}{\beta} \frac{\partial \ln \mathcal{Z}_i}{\partial \mu}$$

$$\mathcal{Z}_i = \sum_{n_i} e^{\beta n_i (\mu - \epsilon_i)}$$

$$\Rightarrow \langle n_i \rangle = \frac{1}{\beta} \frac{\partial \ln \mathcal{Z}_i}{\partial \mu}$$

Fermions

Bosons

$$\mathcal{Z}_i = \sum_{n_i=0,1} e^{\beta n_i (\mu - \epsilon_i)} = 1 + e^{\beta (\mu - \epsilon_i)}$$

geo series

$$\mathcal{Z}_i = \sum_{n_i=0}^{\infty} e^{\beta n_i (\mu - \epsilon_i)} = \frac{1}{1 - e^{\beta (\mu - \epsilon_i)}}$$

together

$$\mathcal{Z}_i = (1 \pm e^{\beta (\mu - \epsilon_i)})^{\pm 1}$$

⊕ fermions
⊖ bosons

converges if $\mu < \epsilon_i$ for $\forall i$
 $\Rightarrow \mu < \epsilon_1 =$ single particle ground state, taken as 0 by convention

$\Rightarrow \mu < 0$ for bosons

chemical potential of a gas of bosons is always negative

$$\langle n_i \rangle = \frac{1}{\beta} \frac{\partial \ln \mathcal{Z}_i}{\partial \mu} = \frac{1}{\beta} \frac{\partial}{\partial \mu} \ln(1 \pm e^{\beta (\mu - \epsilon_i)}) = \frac{1}{\beta} \frac{1}{1 \pm e^{\beta (\mu - \epsilon_i)}} \beta e^{\beta (\mu - \epsilon_i)} \Rightarrow$$

$$\langle n_i \rangle = \frac{1}{e^{\beta (\epsilon_i - \mu)} \pm 1} \leq 1 = f(E)$$

Fermi-Dirac

$$\langle n_i \rangle = \frac{-1}{\beta} \frac{\partial}{\partial \mu} \ln(1 - e^{\beta (\mu - \epsilon_i)}) = \frac{-1}{\beta} \frac{1}{1 - e^{\beta (\mu - \epsilon_i)}} (-\beta e^{\beta (\mu - \epsilon_i)}) =$$

$$\Rightarrow \langle n_i \rangle = \frac{1}{e^{\beta (\epsilon_i - \mu)} - 1} = f(E)$$

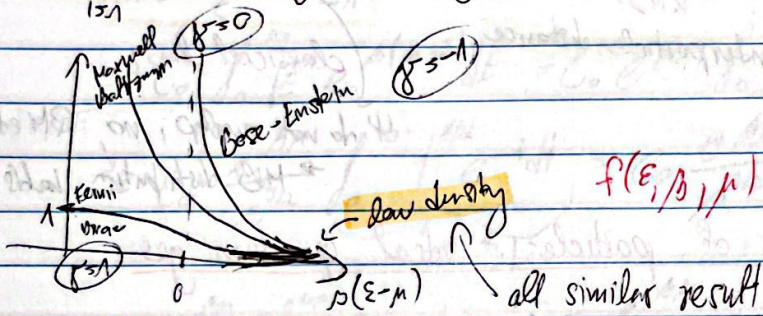
Bose-Einstein

- distributions \forall fixed μ and $T \Rightarrow \epsilon_i \rightarrow E$

\Rightarrow systems in equilibrium w/ heat bath & particle reservoir

$$N = \sum_{i=1}^{\infty} \langle n_i \rangle = \int_0^{\infty} f(E) g(E) dE \quad \text{density of states}$$

$$U = \sum_{i=1}^{\infty} \langle n_i \rangle \epsilon_i = \int_0^{\infty} E f(E) g(E) dE$$



$$f(E, \beta, \mu) = \frac{1}{e^{\beta(E-\mu)} \pm 1}$$

The classical limit

$$\langle n_i \rangle \ll 1, \quad \epsilon_i$$

single particle energies:

$$\epsilon_1 \leq \epsilon_2 \leq \dots \leq \epsilon_i \leq \dots$$

$$e^{\beta(\epsilon_i - \mu)} = e^{-\beta\mu} \leq e^{\beta(\epsilon_i - \mu)} = e^{\beta\epsilon_i} e^{-\beta\mu}$$

$$\epsilon_i \geq 0 \Rightarrow e^{\beta\epsilon_i} \geq 1, \quad \forall i$$

$$\Rightarrow e^{-\beta\mu} \leq e^{\beta(\epsilon_i - \mu)} \quad \forall i$$

$$\langle n_i \rangle = \frac{1}{e^{\beta(\epsilon_i - \mu)} \pm 1}$$

$$\ll 1 \Leftrightarrow e^{\beta(\epsilon_i - \mu)} \gg 1$$

$$\Leftrightarrow e^{-\beta\mu} \gg 1$$

$$\Rightarrow \langle n_i \rangle = \frac{1}{e^{\beta(\epsilon_i - \mu)} \pm 1} \rightarrow e^{-\beta(\mu - \epsilon_i)}$$

classical limit

$$\langle N \rangle = \sum_i \langle n_i \rangle = \sum_i e^{\beta(\mu - \epsilon_i)} = e^{\beta\mu} \sum_i e^{-\beta\epsilon_i} = e^{\beta\mu} Z_1(T, V)$$

(?) When will we have $e^{-\beta\mu} \gg 1$

$$\Rightarrow e^{-\beta\mu} = \frac{Z_1(T, V)}{\langle N \rangle} = \frac{Z_1^{trans} Z_1^{int}}{\langle N \rangle}$$

$$Z_1^{trans} = V \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} = V n_Q$$

$$\lambda_{th} = \left(\frac{1}{n_Q} \right)^{1/3} = \frac{h}{\sqrt{2\pi m k_B T}} = \left(\frac{3}{2\pi} \right)^{1/2} \lambda_{dB}$$

$$\lambda_{th} \approx \lambda_{dB}$$

$$Z_1^{int} = \sum_{\alpha=1}^{\infty} e^{-\beta \epsilon_{\alpha}^{int}} = 1 + \sum_{\alpha=2}^{\infty} e^{-\beta \epsilon_{\alpha}^{int}} \geq 1$$

$$e^{-\beta\mu} \gg 1 \text{ if } \frac{V}{\langle N \rangle} \frac{1}{\lambda_{th}^3} \gg 1$$

$$\Rightarrow e^{-\beta\mu} = \frac{V}{\langle N \rangle} \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} Z_1^{int} \geq \frac{V}{\langle N \rangle} \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} = \frac{V}{\langle N \rangle} \lambda_{th}^{-3}$$

$$\frac{V}{N} \frac{1}{\lambda_{th}^3} \gg 1 \Rightarrow \left(\frac{p}{\hbar}\right)^3 \gg \left(\frac{p}{\hbar_{th}}\right)^3 \gg 1$$

$l = \left(\frac{V}{N}\right)^{1/3}$
interparticle distance

$$p \gg \hbar_{th} \approx \hbar_{th}$$

classical limit

ψ do not overlap, no QM effects,
→ MB distribution holds

Numbers of particles - ideal quantum gas

$$N = \sum_i \langle n_i \rangle = \sum_i \frac{1}{e^{\beta(\epsilon_i - \mu)} \pm 1}$$

integral

$$N = \int_0^{\infty} g(E) \frac{1}{e^{\beta(E - \mu)} \pm 1} dE$$

$E = \frac{p^2}{2m}$

$$g(p) dp = \frac{V}{h^3} 4\pi p^2 dp$$

$$g(E) dE = \frac{V}{h^3} 2\pi(2m)^{3/2} \sqrt{E} dE$$

$$N = \frac{V}{h^3} 2\pi(2m)^{3/2} \int_0^{\infty} \frac{\sqrt{E} dE}{e^{\beta(E - \mu)} \pm 1}$$

$g = 2s + 1$ = # possible spin states

Internal energy - ideal quantum gas

$$U = \sum_i \langle n_i \rangle \epsilon_i = \sum_i \frac{\epsilon_i}{e^{\beta(\epsilon_i - \mu)} \pm 1}$$

sum
integral

$$U = \int g(E) \frac{E}{e^{\beta(E - \mu)} \pm 1} dE \Rightarrow$$

$$U = \frac{V}{h^3} 2\pi(2m)^{3/2} \int_0^{\infty} \frac{E^{3/2}}{e^{\beta(E - \mu)} \pm 1} dE$$

Grand potential - ideal quantum gas

$$\Phi_G = -\frac{2}{3} U$$

$$\Phi_G = -pV$$

$$\Phi_G = -\frac{1}{\beta} \ln \mathcal{Z} = -\frac{1}{\beta} \ln \left(\prod_i \mathcal{Z}_i \right) = -\frac{1}{\beta} \sum_i \ln \mathcal{Z}_i$$

$$\Rightarrow \Phi_G = \mp \frac{1}{\beta} \sum_i \ln(1 \pm e^{\beta(\mu - \epsilon_i)})$$

⊕ fermions
⊖ bosons

integral

$$\Phi_G = \mp \frac{1}{\beta} C_0 \int_0^{\infty} \ln(1 \pm e^{\beta(\mu - E)}) \sqrt{E} dE$$

by parts

$$= \mp \frac{2}{3} \frac{C_0}{\beta} E^{3/2} \ln(1 \pm e^{\beta(\mu - E)}) \Big|_0^{\infty} + \frac{2}{3} \frac{C_0}{\beta} \int_0^{\infty} E^{3/2} d(\ln(1 \pm e^{\beta(\mu - E)}))$$

$E = \infty: e^{-\beta E} \rightarrow 0$
 $E = 0: E \rightarrow 0$

$$= \mp \frac{2}{3} \frac{C_0}{\beta} \int_0^{\infty} \frac{E^{3/2} \pm e^{\beta(\mu - E)} (-\beta) E^{3/2}}{1 \pm e^{\beta(\mu - E)}} dE = -\frac{2}{3} C_0 \int_0^{\infty} \frac{E^{3/2}}{e^{\beta(E - \mu)} \pm 1} dE$$

Using classical limit

$$e^{-\beta\mu} \gg 1 \Rightarrow e^{\beta(E-\mu)} \gg 1$$

$$N = C_0 \int_0^\infty \frac{\sqrt{E}}{e^{\beta(E-\mu)} \pm 1} dE$$

$$\begin{aligned} &\Rightarrow C_0 \int_0^\infty \sqrt{E} e^{\beta(\mu-E)} dE = C_0 e^{\beta\mu} \int_0^\infty \sqrt{E} e^{-\beta E} dE \\ &\stackrel{t = \beta E}{\Rightarrow} C_0 \frac{e^{\beta\mu}}{\beta^{3/2}} \int_0^\infty t^{1/2} e^{-t} dt = C_0 \frac{e^{\beta\mu}}{\beta^{3/2}} \Gamma\left(\frac{3}{2}\right) \end{aligned}$$

$$U = C_0 \frac{e^{\beta\mu}}{\beta^{5/2}} \int_0^\infty t^{3/2} e^{-t} dt = C_0 \frac{e^{\beta\mu}}{\beta^{5/2}} \Gamma\left(\frac{5}{2}\right) = C_0 \frac{e^{\beta\mu}}{\beta^{5/2}} \frac{3}{2} \Gamma\left(\frac{3}{2}\right)$$

$$\Rightarrow \left. \begin{aligned} U &= \frac{3}{2} \frac{N}{\beta} = \frac{3}{2} N k_B T \\ \mathcal{F}_0 &= -\frac{2}{3} U = -PV \end{aligned} \right\} PV = Nk_B T \checkmark$$

$$\Gamma(n) = \int_0^\infty t^{n-1} e^{-t} dt$$

$$\Gamma(n+1) = n \Gamma(n)$$

Fermions & Bosons

$$N = C_0 \int_0^\infty \frac{E^{1/2}}{e^{\beta(E-\mu)} \pm 1} dE = C_0 \int_0^\infty \frac{E^{1/2}}{z^{-1} e^{\beta E} \pm 1} dE =$$

$$\stackrel{x = \frac{E}{k_B T}}{=} C_0 (k_B T)^{3/2} \int_0^\infty \frac{x^{1/2}}{z^{-1} e^{x} \pm 1} dx \quad \begin{matrix} z = e^{\beta\mu} \\ \text{fugacity} \end{matrix}$$

$$\Rightarrow \left. \begin{aligned} \text{Fermi-Dirac integral} & f_n(z) = \int_0^\infty \frac{x^n}{z^{-1} e^x \pm 1} dx \\ \text{Bose-Einstein integral} & g_n(z) = \int_0^\infty \frac{x^n}{z^{-1} e^x - 1} dx \end{aligned} \right\}$$

$$\begin{aligned} g_n(z) &= \int_0^\infty \frac{x^n}{z^{-1} e^x - 1} dx = \int_0^\infty x^n \frac{z e^{-x}}{1 - z e^{-x}} dx = \int_0^\infty x^n \left\{ z e^{-x} \sum_{m=0}^{\infty} (z e^{-x})^m \right\} dx \\ &= \sum_{m=0}^{\infty} z^{m+1} \int_0^\infty x^n e^{-(m+1)x} dx = \sum_{m=0}^{\infty} \frac{z^{m+1}}{(m+1)^{n+1}} \int_0^\infty y^n e^{-y} dy \\ &= \Gamma(n+1) \sum_{m=0}^{\infty} \frac{z^{m+1}}{(m+1)^{n+1}} = \Gamma(n+1) \sum_{k=1}^{\infty} \frac{z^k}{k^{n+1}} = \Gamma(n+1) \text{Li}_{n+1}(z) \\ &= \text{Li}_{n+1}(z) \end{aligned}$$

simi logic for $f_n(z)$:

$$\left. \begin{aligned} f_n(z) &= -\Gamma(n+1) \text{Li}_{n+1}(-z) \\ g_n(z) &= \Gamma(n+1) \text{Li}_{n+1}(z) \end{aligned} \right\}$$

= polylogarithm

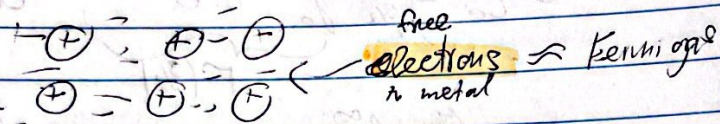
$$N = C_0 (k_B T)^{3/2} \int_0^\infty \frac{x^{1/2}}{e^{-x} + 1} dx \quad n=5/2 \quad \text{FD BE}$$

$$U = C_0 (k_B T)^{5/2} \int_0^\infty \frac{x^{3/2}}{e^{-x} + 1} dx = \frac{3}{2} \Gamma\left(\frac{5}{2}\right) \text{Li}_5\left(-\frac{1}{z}\right)$$

$$\Rightarrow U = \frac{\Gamma(5/2)}{\Gamma(3/2)} N k_B T \frac{\text{Li}_{5/2}(-z)}{\text{Li}_{3/2}(-z)} = \frac{3}{2} N k_B T \frac{\text{Li}_{5/2}(-z)}{\text{Li}_{3/2}(-z)}$$

classical quantum effects

Fermi gas
metal



$$N = C_0 \int_0^\mu \frac{E^2}{e^{E-\mu} + 1} dE$$

$$C_0 = \frac{2m^3}{h^3} (2\pi)^{3/2} = \frac{4\pi V}{h^3} (2m)^{3/2}$$

$$\Rightarrow N = \frac{4\pi V}{h^3} (2m)^{3/2} \int_0^\mu \frac{E^2 dE}{e^{E-\mu} + 1}$$

2s+1 = 2 for electrons

↳ if N, V are known $\Rightarrow \mu = \mu(T)$ by solving eq.

↳ define Fermi energy $E_F = \mu(T=0)$

claim: $E_F > 0$

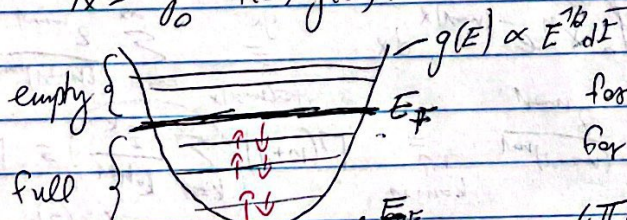
↳ skip proof - see lecture 15

$$f(E) = \frac{1}{e^{\beta(E-\mu)} + 1} \xrightarrow{T \rightarrow 0} \frac{1}{e^{\beta(E-E_F)} + 1} = \frac{1}{e^{\frac{E-E_F}{k_B T}} + 1}$$

if $E > E_F \Rightarrow e^{\frac{E-E_F}{k_B T}} \rightarrow e^\infty \Rightarrow f(E) = 0$
 if $E < E_F \Rightarrow e^{\frac{E-E_F}{k_B T}} \rightarrow e^{-\infty} \Rightarrow f(E) = 1$

Electrons (spin-1/2 fermions) at $T=0$

$$N = \int_0^\infty f(E) g(E) dE$$



for $E > E_F$: $f(E) = 0$

for $E < E_F$: $f(E) = 1$

$$\Rightarrow N = \int_0^{E_F} g(E) dE = \frac{4\pi V}{h^3} (2m)^{3/2} \int_0^{E_F} E^{1/2} dE = \frac{4\pi V}{h^3} (2m)^{3/2} \frac{2}{3} E_F^{3/2}$$

$$\Rightarrow E_F = \frac{h^2}{2m} \left(\frac{3}{8\pi} \frac{N}{V} \right)^{2/3} = \frac{2}{3} k_B T_F$$

$$U = \int_0^{\infty} E f(E) g(E) dE \approx \int_0^{E_F} E g(E) dE = \frac{4\pi V}{h^3} (2m)^{3/2} \int_0^{E_F} E^{3/2} dE$$

$$U = \frac{4\pi V}{h^3} (2m)^{3/2} \frac{2}{5} E_F^{5/2} \stackrel{\text{at } T=0}{=} \frac{3}{5} N E_F = \frac{3}{5} N k_B T_F$$

$$E_F = \frac{\hbar^2}{2m} \left(\frac{6\pi^2}{2s+1} \frac{N}{V} \right)^{2/3} = k_B T_F = \frac{\hbar^2}{2m} \left(\frac{3}{4\pi(2s+1)} \frac{N}{V} \right)^{2/3}$$

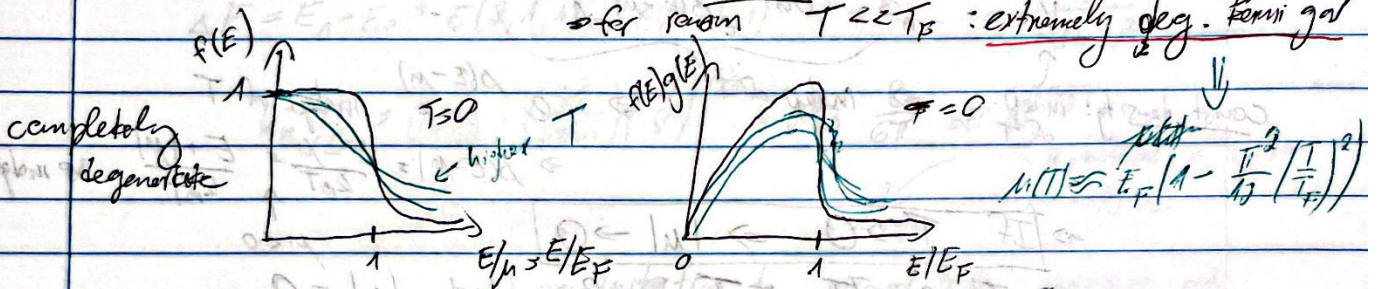
$$E_F = \frac{\hbar^2}{2m} \left(\frac{3}{8\pi} \frac{N}{V} \right)^{2/3} = k_B T_F \quad \text{for electrons } s=1/2$$

For metals: $\frac{N}{V} = 10^{22} - 10^{23} \text{ electrons/cm}^3$

$$\Rightarrow E_F \approx 5 \text{ eV}$$

$$\Rightarrow T_F \approx 5 \cdot 10^4 \text{ K} \quad \text{- much higher than room } T$$

\Rightarrow for room $T \ll T_F$: extremely deg. Fermi gas



Electronic heat capacity

$$C_V = \underbrace{\frac{\pi^2}{2} R \frac{T}{T_F}}_{\text{electrons}} + \underbrace{\frac{12}{5} \pi^4 R \left(\frac{T}{\Theta_0} \right)^3}_{\text{lattice vibrations}} \quad \text{per mole}$$

- start with degenerate Fermi gas
- raise temp. to $T \ll T_F \Rightarrow e^-$ gain energy $\sim k_B T$
- only electrons close to Fermi energy can make the jump to higher levels

electrons promoted to higher level: $\Delta N \approx f(E_F) g(E_F) k_B T \approx 1$

$$g(E_F) = \frac{4\pi V}{h^3} (2m)^{3/2} \sqrt{E_F}$$

$$N = \frac{4\pi V}{h^3} (2m)^{3/2} \frac{2}{3} E_F^{3/2}$$

$$\Delta N \approx \frac{3}{2} \frac{N}{E_F} k_B T = \frac{3}{2} N \frac{T}{T_F}$$

$$\Delta E \approx \Delta N k_B T = \frac{3}{2} N k_B \frac{T^2}{T_F}$$

$$T_F = \frac{E_F}{k_B} = \frac{\hbar^2}{2mk_B} \left(\frac{3\pi^2 N}{V} \right)^{2/3}$$

$$\Theta_D = \frac{\hbar v_D}{k_B} = \frac{\hbar}{k_B} \left(\frac{6\pi^2 N}{V} \right)^{1/3} v_D$$

$$C_V = \frac{\partial E}{\partial T} \approx 3N k_B \frac{T}{T_F} \approx 3nR \frac{T}{T_F}$$

more sophisticated treatment: $C_V = \frac{\pi^2}{2} nR \frac{T}{T_F}$

• at room temp. $\frac{T}{T_F}$ is small \Rightarrow electron contribution very small compared to $3R$ of lattice

Bose-Einstein condensation

- ideal gas of bosons of spin 0
- BE occupation number

$$\langle n_i \rangle = \frac{1}{e^{\beta(E_i - \mu)} - 1}$$

• total # bosons: $N = \sum \langle n_i \rangle$

$$g(E)dE = \frac{V}{h^3} 2\pi(2m)^{3/2} E^{1/2} dE$$

$$N = \frac{2\pi V (2m)^{3/2}}{h^3} \int_0^\infty \frac{E^{1/2}}{e^{\beta(E - \mu)} - 1} dE$$

$$\frac{N}{V} = \frac{2\pi(2m)^{3/2}}{h^3} \int_0^\infty \frac{E^{1/2}}{e^{\beta(E - \mu)} - 1} dE$$

const. density: indep. of T \Rightarrow indep. of T $\Rightarrow e^{\beta(E - \mu)}$ indep. of T

$$\Rightarrow \beta(E - \mu) = \frac{E - \mu}{k_B T} = \frac{E + |\mu|}{k_B T} \text{ is indep. of T}$$

$\mu < 0$

$$\Rightarrow \text{if } T \rightarrow 0 \Rightarrow |\mu| \rightarrow 0$$

\Rightarrow There is $T = T_c$ for which $|\mu| = 0$

\Rightarrow thus it seems that the gas cannot be cooled below T_c

Condensation temperature

at T_c : $\frac{N}{V} = \frac{2\pi(2m)^{3/2}}{h^3} (k_B T_c)^{3/2} \int_0^\infty \frac{\sqrt{x}}{e^x - 1} dx$

$$x = \frac{E}{k_B T_c}$$

$$\Rightarrow \frac{N}{V} = 2.612 \cdot \left(\frac{2\pi m k_B T_c}{h^2} \right)^{3/2} \cdot \frac{N}{2}$$

\leftarrow defined in T_c

When moving from grand sum \Rightarrow integral,

we assumed $E_0 = 0 \Rightarrow$ no contribution to integral

\Rightarrow causing trouble now because

there's many particles in ground state at low T

$$N = \sum \frac{1}{e^{\beta(E_i - \mu)} - 1} = N_0 + N(E > 0) = \frac{1}{e^{-\beta\mu} - 1} + \frac{2\pi V (2m)^{3/2}}{h^3} \int_0^\infty \frac{E^{1/2}}{e^{\beta(E - \mu)} - 1} dE$$

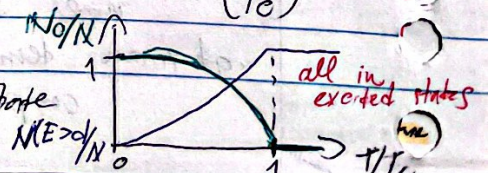
particles in ground state with $E = 0$

$$N_0 = N - N(E > 0)$$

$$\frac{N_0}{N} = 1 - \frac{N(E > 0)}{N} = 1 - \left(\frac{T}{T_c} \right)^{3/2}$$

$$N(E > 0) = N \left(\frac{T}{T_c} \right)^{3/2}$$

as earlier



BEC in rubidium vapor

10000 Rb atoms in a box of $\sim 10 \mu\text{m}$

$$\frac{N}{V} \sim \frac{10^4}{(10^{-5})^3} \rightarrow 10^{19} \text{ m}^{-3}$$

interparticle spacing: $\rho \sim \frac{1}{(10^{19})^{1/3}} \sim 5 \cdot 10^{-7} \text{ m}$ } extremely
 atomic radius Rb: $r_0 \approx 2.4 \cdot 10^{-8} \text{ m}$ } dilute
 gas

$$\frac{N}{V} = 2.61 \left(\frac{2\pi m k_B T_c}{h^2} \right)^{3/2} \quad m \approx 1.4 \cdot 10^{-25} \text{ kg} \quad \Rightarrow T_c = 8.2 \cdot 10^{-8} \text{ K}$$

• particles in a box:

$$E = \frac{p^2}{2m} = \frac{\hbar^2 (k_x^2 + k_y^2 + k_z^2)}{2mL^2}$$

$$\Delta E = E_1 - E_0 = E(2,1,1) - E(1,1,1) = \dots = \frac{3\hbar^2}{8mL^2}$$

$$k_B T_c = 0.53 \cdot \left(\frac{h^2}{2\pi m} \right) \left(\frac{N}{V} \right)^{2/3} = 0.53 \frac{8}{6\pi} \frac{\hbar^2}{2m} \frac{N}{V}^{2/3}$$

particles in system
 large

$$k_B T_c \gg \Delta E$$

\Rightarrow condensation to ground state occurs at thermal energies much larger than the distance between the ground state

and first excited state

$$V_Q = (\lambda_{th})^3 = \frac{h^3}{(2\pi m k_B T_c)^{3/2}} \approx 2.7 \cdot 10^{-19} \text{ m}^3$$

Physical volume: $V_{phys} \approx \frac{(10^{-5})^3}{10000} \approx 10^{-19} \text{ m}^3$

$$\Rightarrow V_Q \approx V_{phys}$$

\hookrightarrow although the gas is very dilute, quantum aspects become important because wavelengths of particles overlap

For $T < T_c$ only consider particles not in ground state

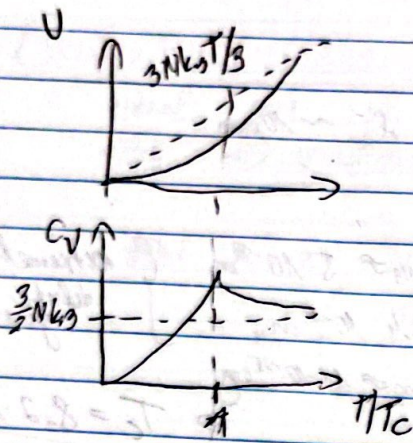
because $E_0 = 0$

$$U = \frac{2\pi V}{h^3} (2m)^{3/2} \int_0^\infty \frac{E^{3/2}}{e^{(E/k_B T)} - 1} dE \approx \frac{2\pi V}{h^3} (2m)^{3/2} \int_0^\infty \frac{E^{3/2}}{e^{BE-1}} dE$$

$$= \frac{2\pi V}{h^3} (2m)^{3/2} (k_B T)^{5/2} \int_0^\infty \frac{x^{3/2}}{e^x - 1} dx \quad \left[\frac{T_c}{T} \approx 150 \right] \approx 2.01 \cdot V \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} k_B T$$

$$\Rightarrow U = 0.77 N k_B \left(\frac{T}{T_c} \right)^{5/2} \quad \frac{N}{V} = ? \text{ as above}$$

$$C_v = \frac{\partial U}{\partial T} = \frac{5}{2} 0.77 N k_B \left(\frac{T}{T_c} \right)^{3/2} = 1.92 N k_B \left(\frac{T}{T_c} \right)^{3/2}$$



Why grand state problem matters for bosons but not fermions?

- only 1 or 2 fermions in grand state \Rightarrow negligible
- but at low T , many bosons in grand state